

FNAL/ANL Processing and Analysis of PreCam Data

(with a bit about hardware too)

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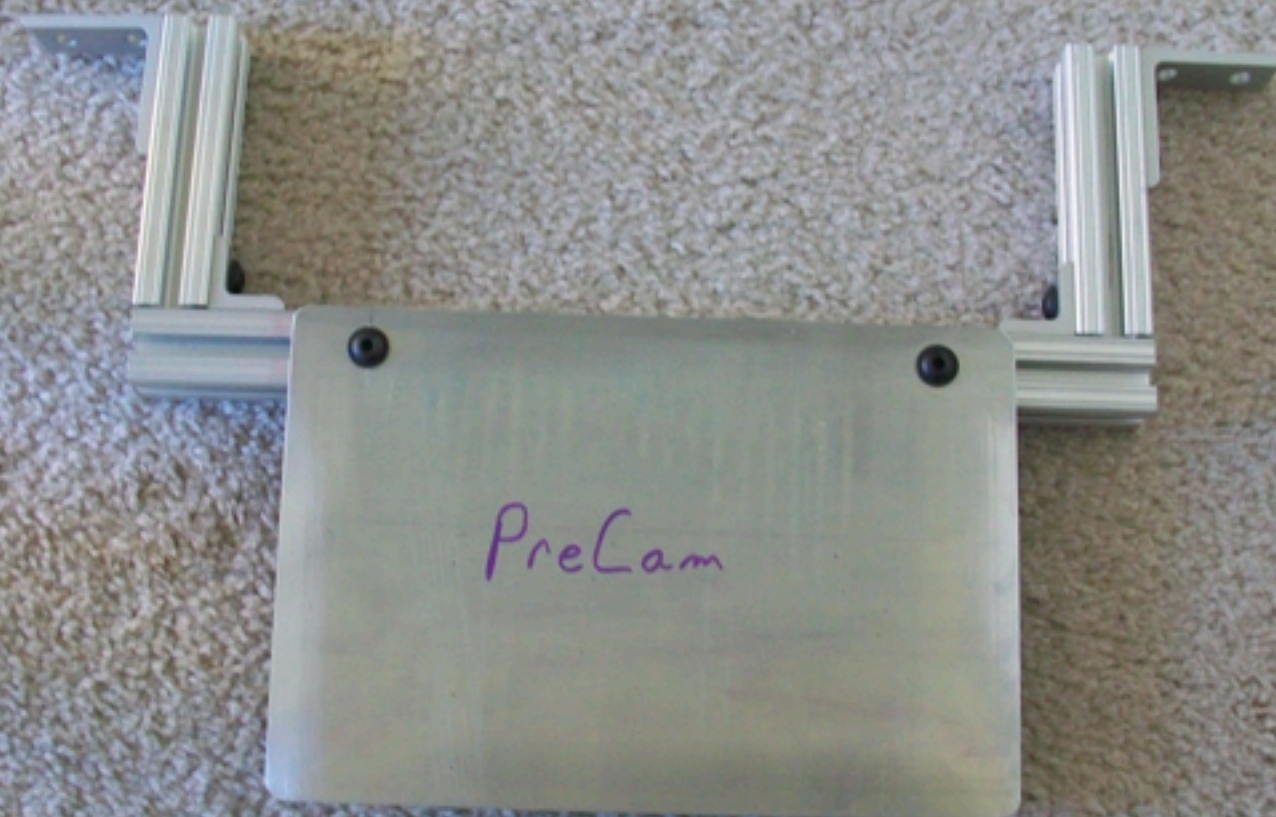
DES Collaboration Meeting
University of Pennsylvania
October 8–11, 2011

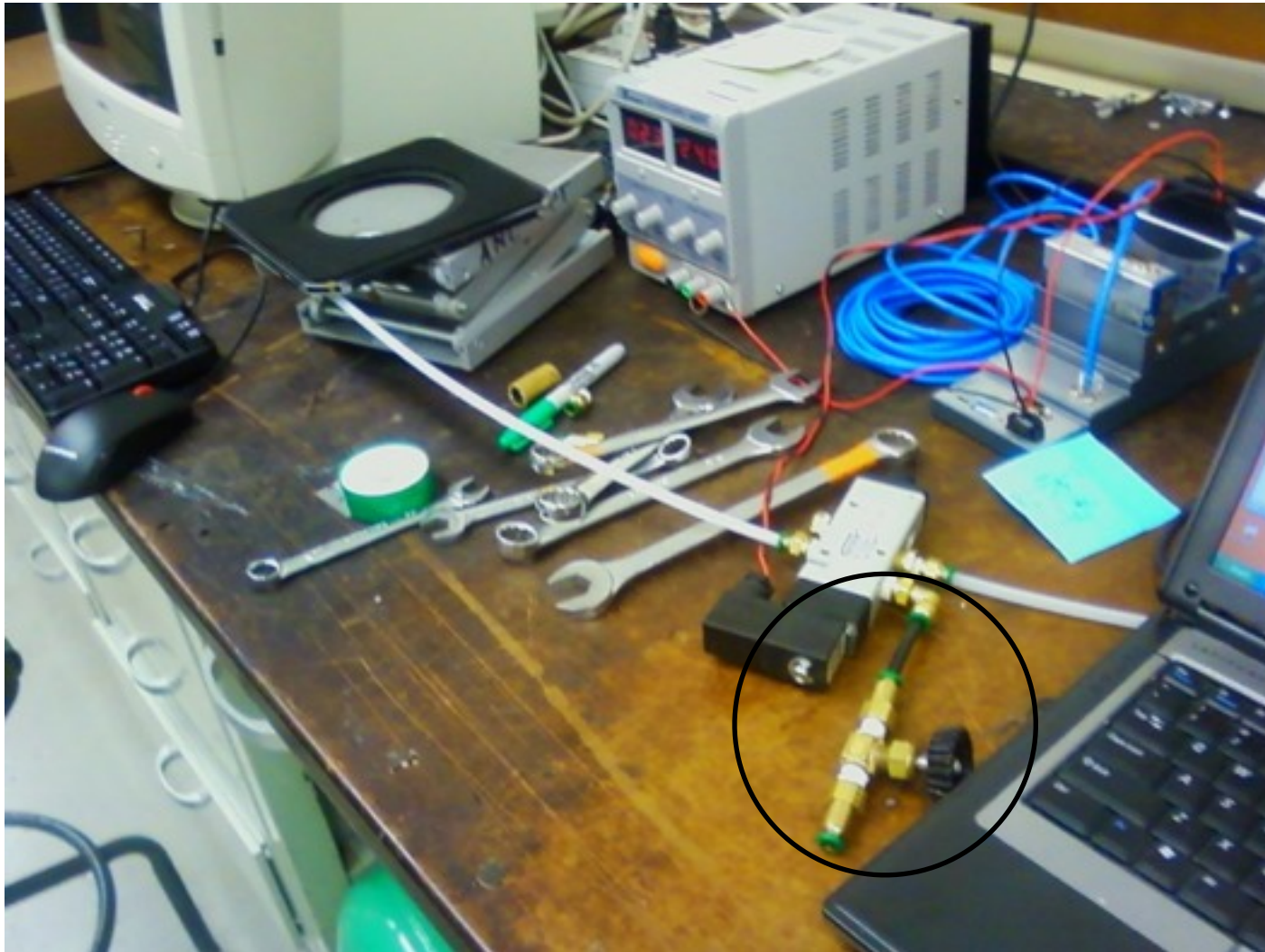


In this talk...

- Hardware
 - VIB Cables “fixed”, strain relief shipped
 - Shutter: >45k actuations in 6 positions, no failures (~10–20% of all actuations directly observed), thanks to slow-release valve...
- Overview of processing: v1, v2, v3
- FNAL/ANL shell/py scripts:
 - Bias, flat, dark, etc. (standard processing)
 - Streaking correction
 - Stuck shutter image identification
 - shutter timing & illumination map corrections
- ANL root scripts incorporating star flat corrections, data selection criteria--particularly FWHM, stellarity, magnitude (error)
- Results: Astrometry and Photometry for USNO, South, SDSS
- Other cool products: Montage, SNe in PreCam
- To Do: Alternate photometric solutions (DAOPHOT, PSFEx), any additional errors (SDSS reCal?), FNAL/ANL scripts to Brazil pipeline
- PUBLISH!

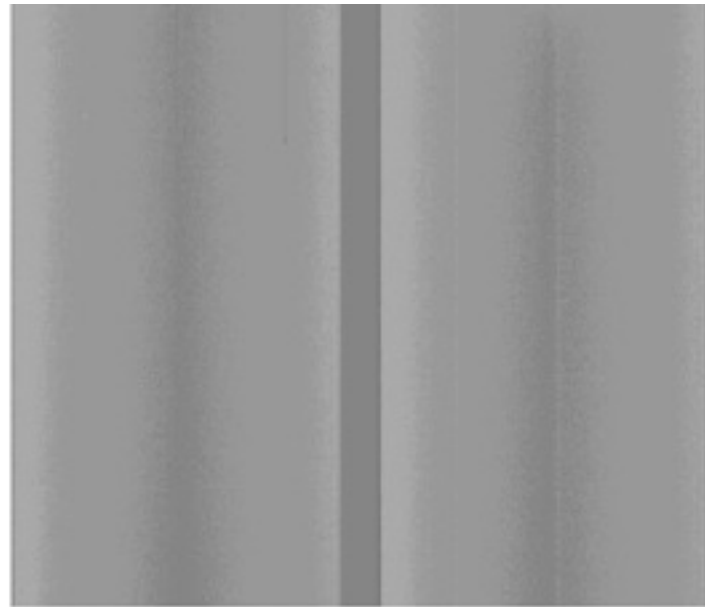




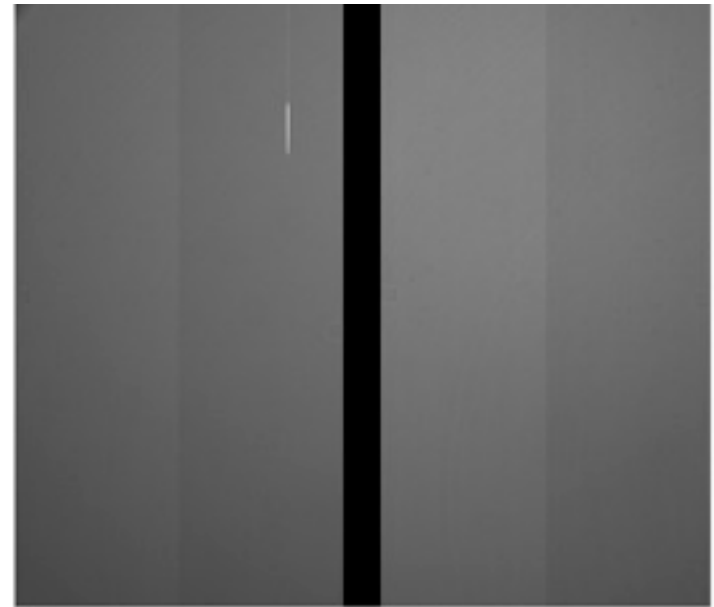


FNAL v1, v2, v3 processing (doc-db#5872)

- Fermilab developed a processing pipeline consisting primarily of shell and py scripts for bias subtraction, flat-field corrections, etc.
- Each iteration added functionality--crucial improvements include banding/streaking removal, astrometry
- since v3, shutter timing maps and stuck shutter identification and correction algorithms have also been developed

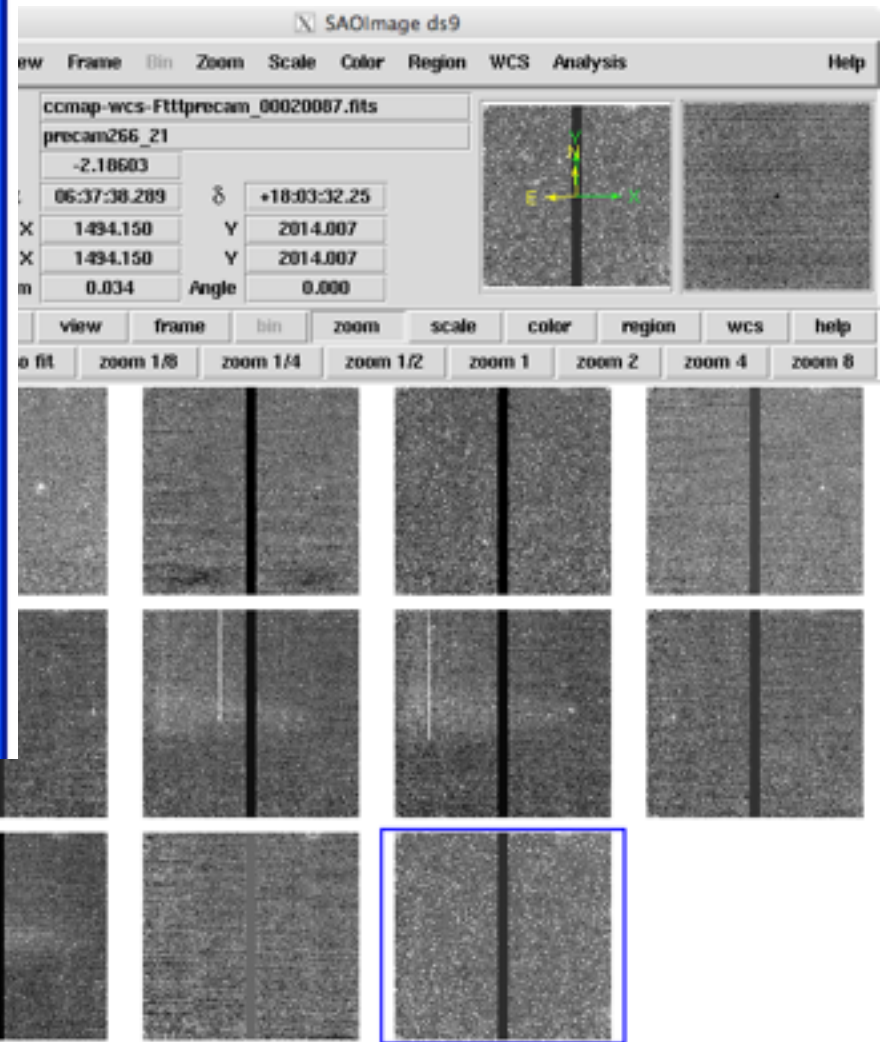
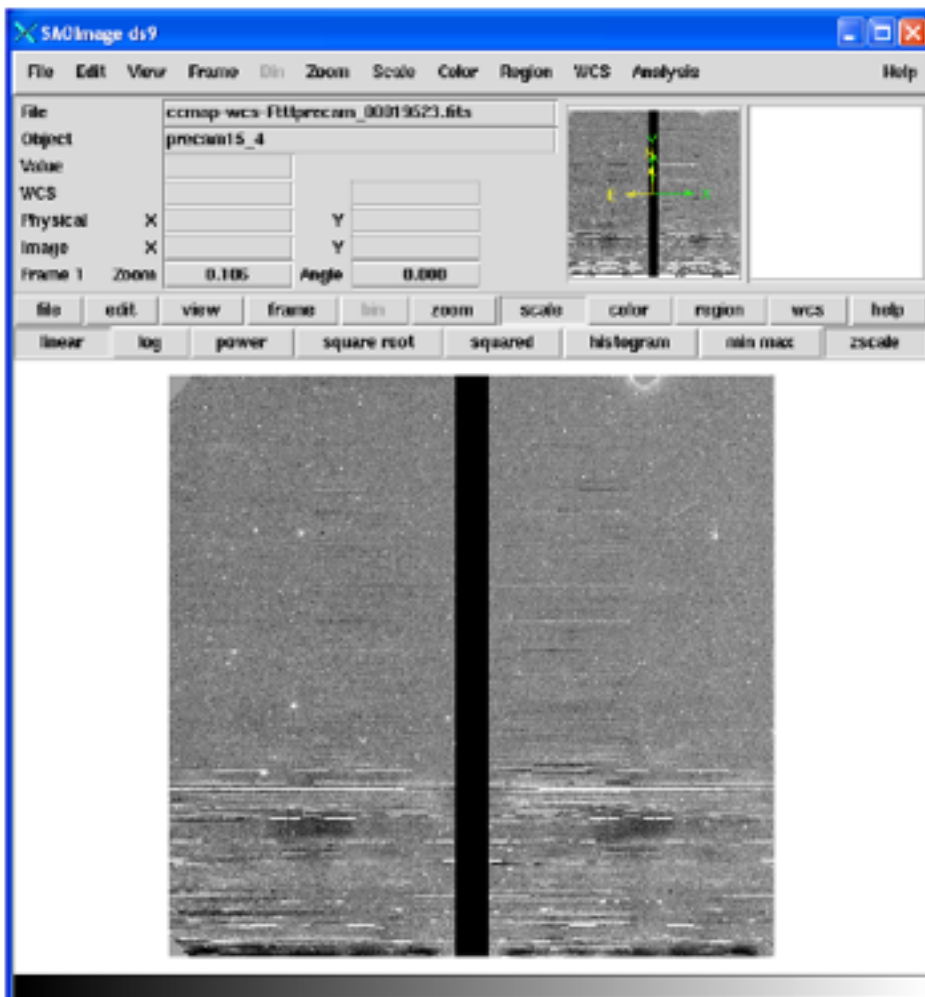


Bias

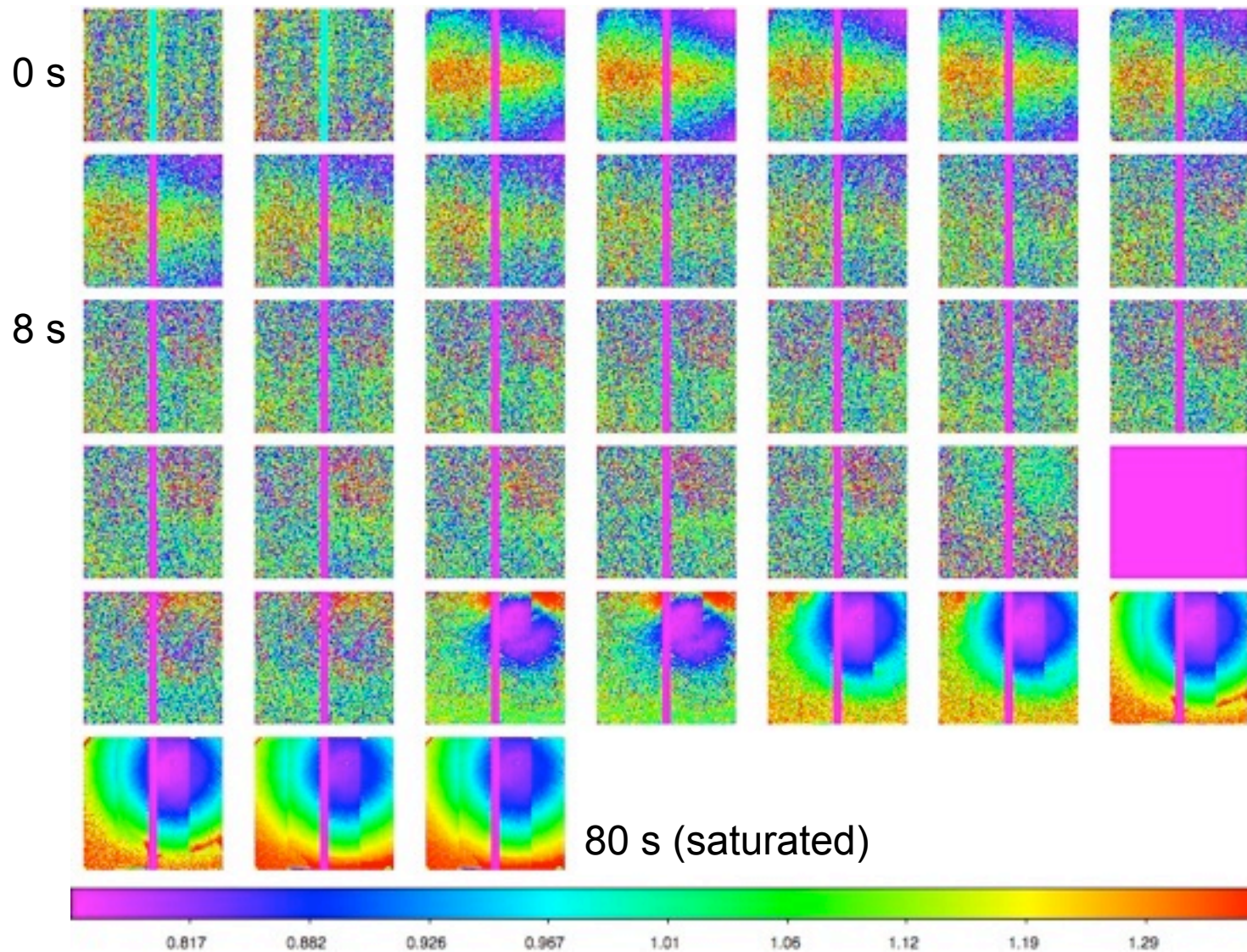


Flat





From Photon Transfer Curve data: Shutter Timing effects are negligible beyond ~8s.



FNAL v3+ Processing Status

- comprehensive bad file list compiled
- corrections for streaking, stuck shutter, shutter timing, etc. applied
- (nearly) all files completely processed, *.cat files produced

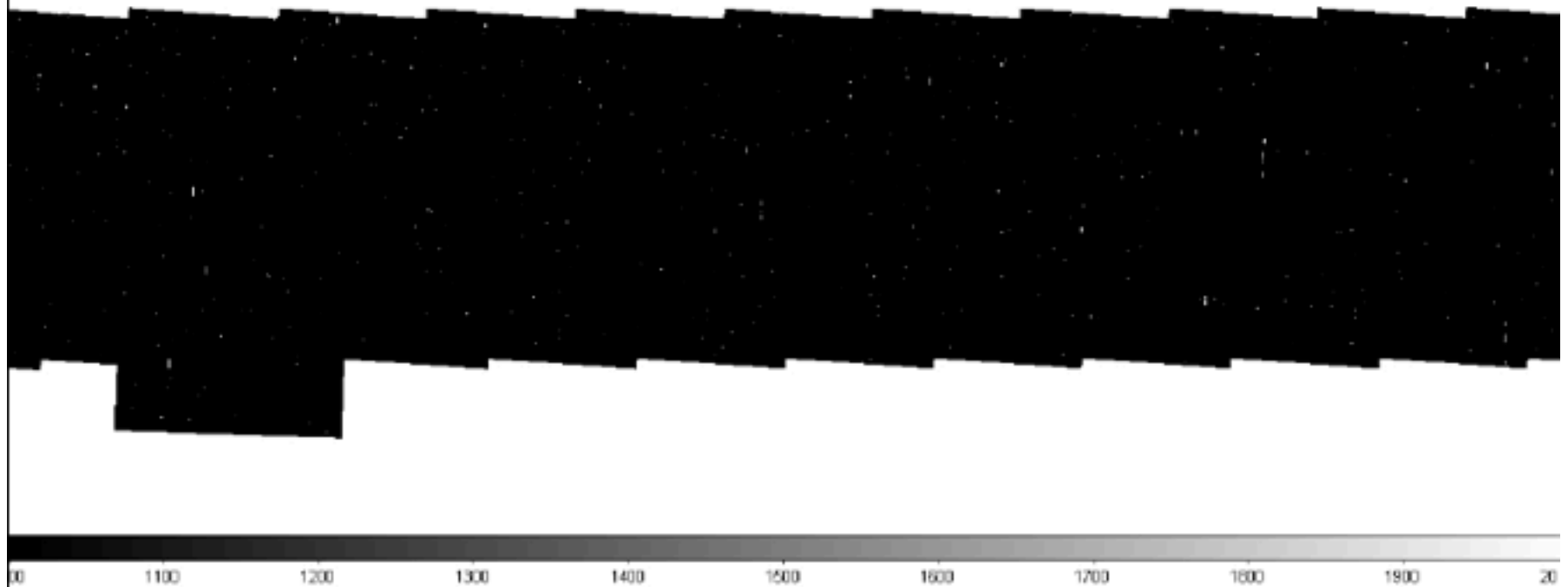
Argonne has been hard at work in parallel...

- (Hardware: shutter improvements, VIB Cable strain relief)
- Montage!
- Additions to processing including independent stuck shutter identification, as well as illumination correction (secondary flat field due to shutter or other instrumental effects, and dipoles/artifacts from streaking correction algorithm) --> new .cat files
- Analysis tools (primarily root scripts) developed and applied to Golden Nights

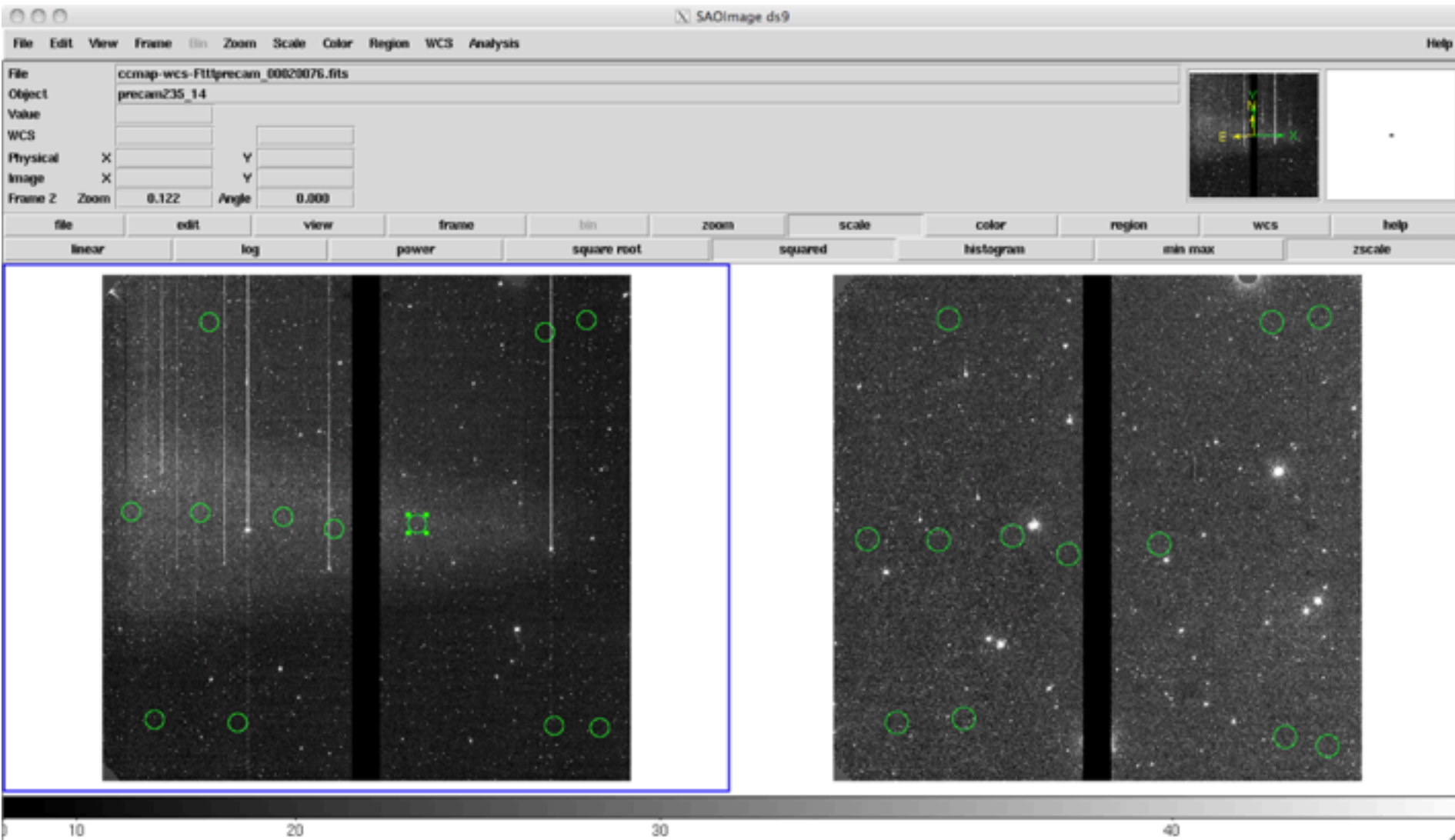


Montage!

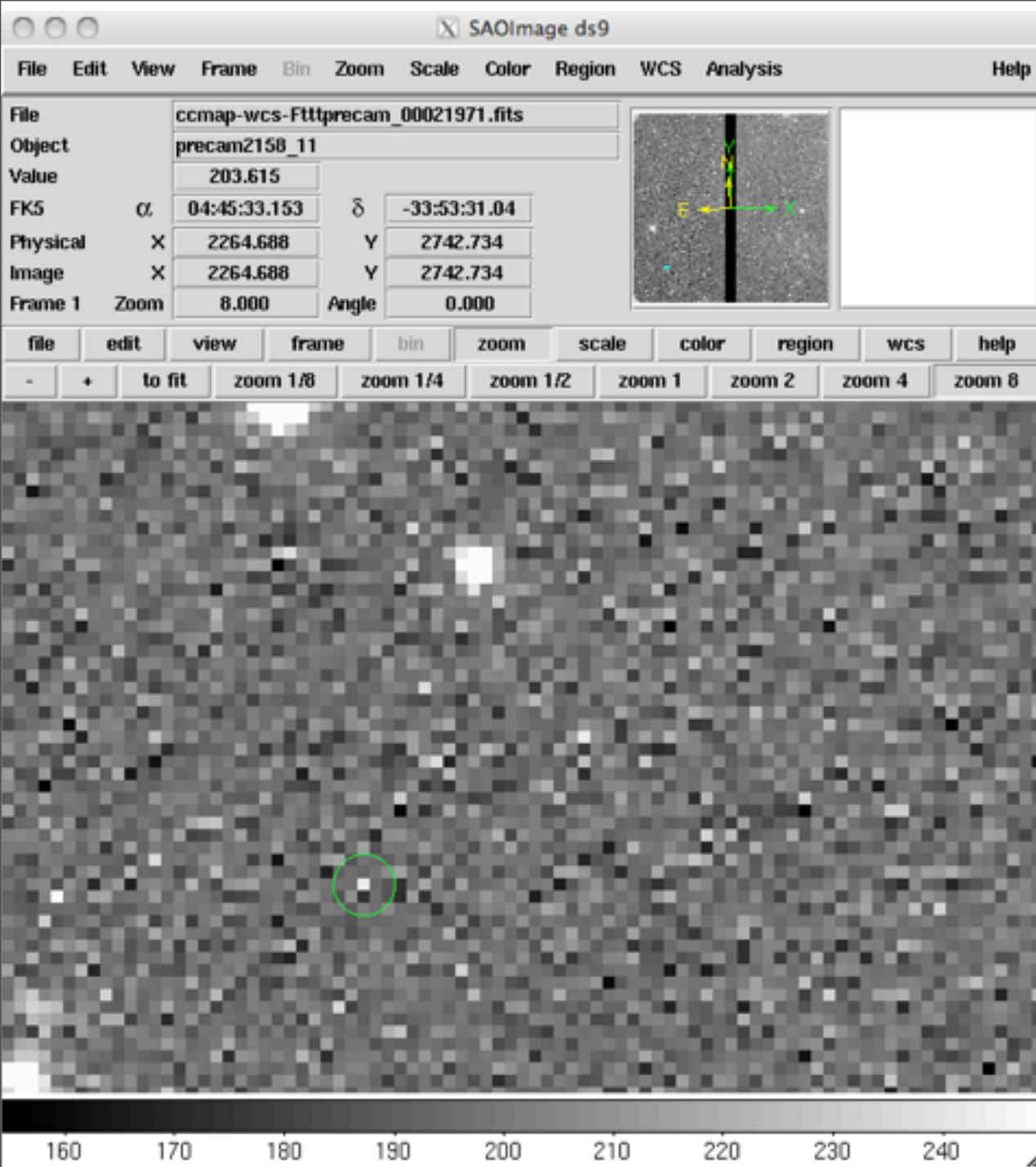
53-image i-band only PreCam mosaic of Stripe 82, with scale chosen to hide CCD gaps. Processing took 1 minute. Also used Montage feature of pixel-averaging to shrink combined FITS file by x20.



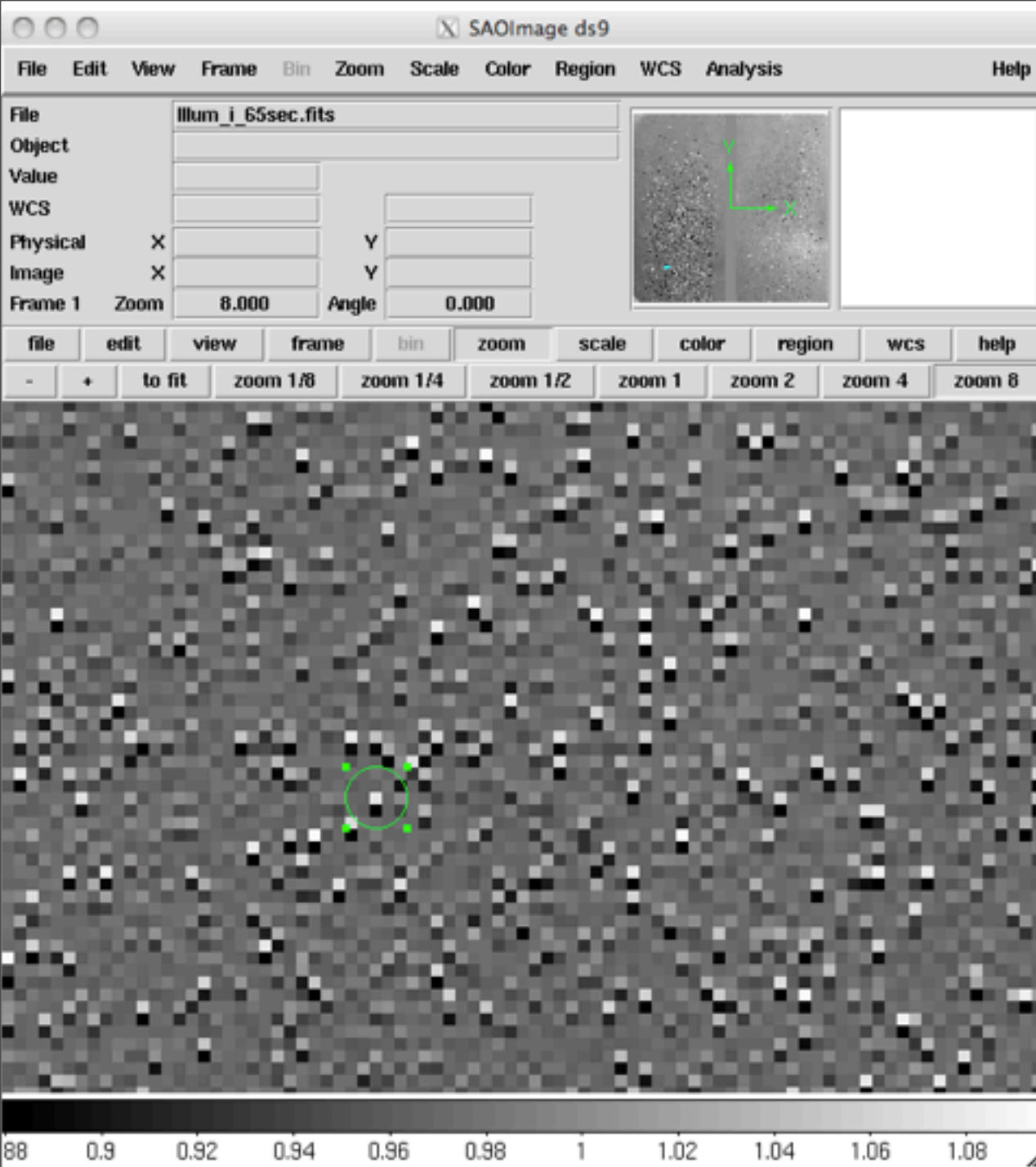
Shutter Stuck? Test



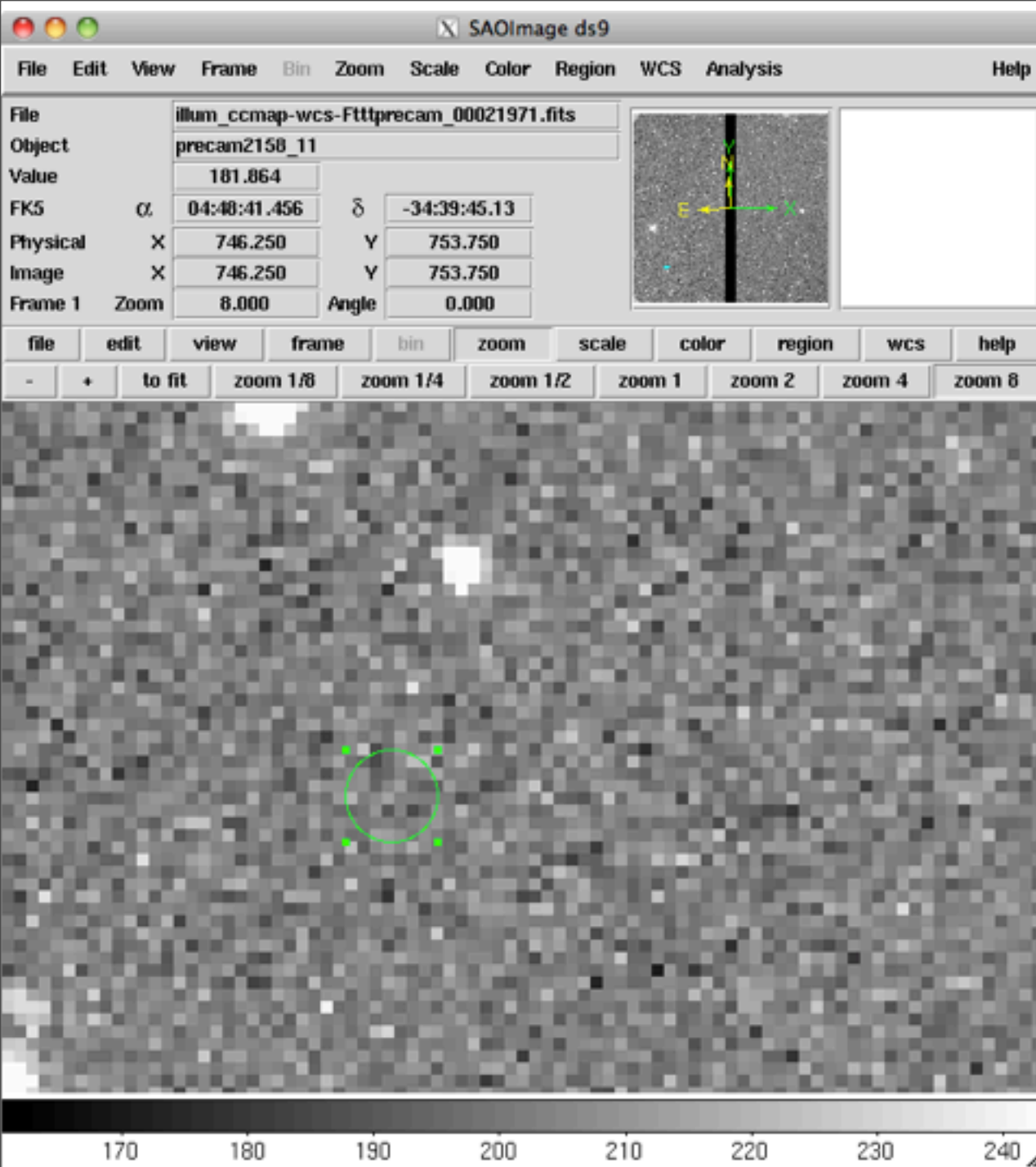
Moving Shutter region to centerline, Weighting BG points improves separation.
NO false positives, only 2 false negatives (detectable by eye)



ccmap-*.fits
Highlighted Dipole
at x,y = 756, 740:
Top = 242.3
Bottom = 174.3
+/- 34 from mean
+/- 16% from mean



Illumination Map
same x,y
Top = 1.079
Bottom = 0.838



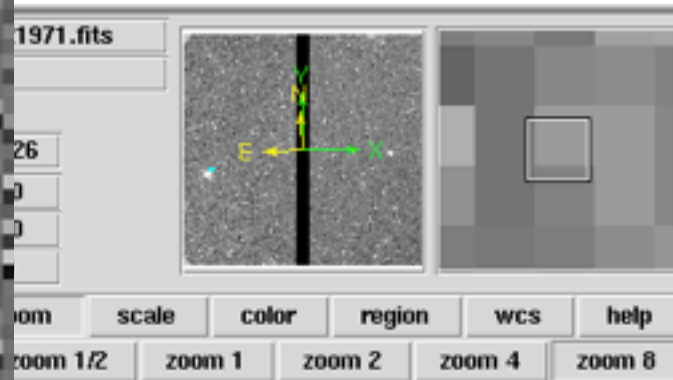
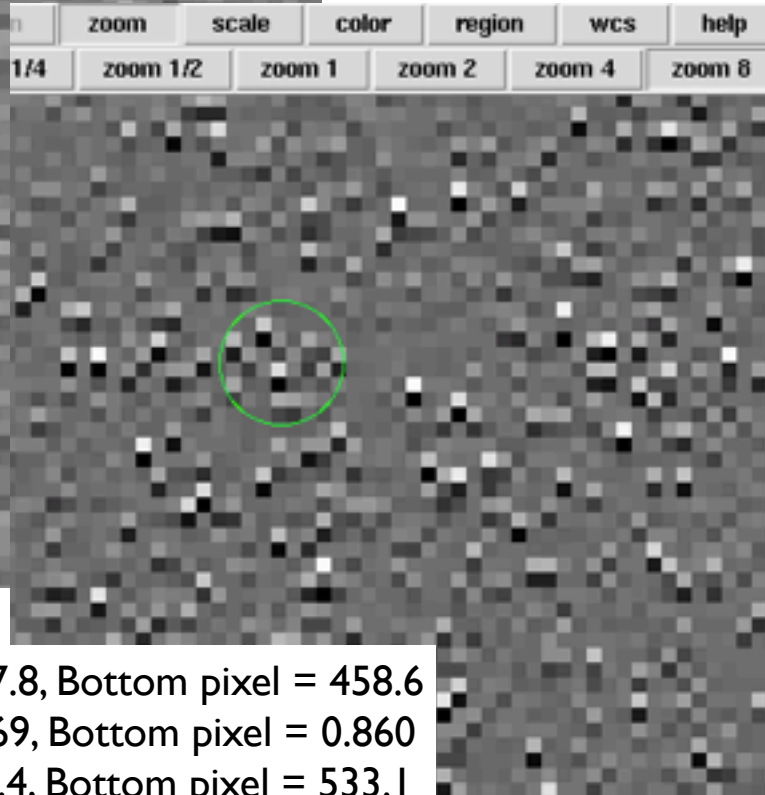
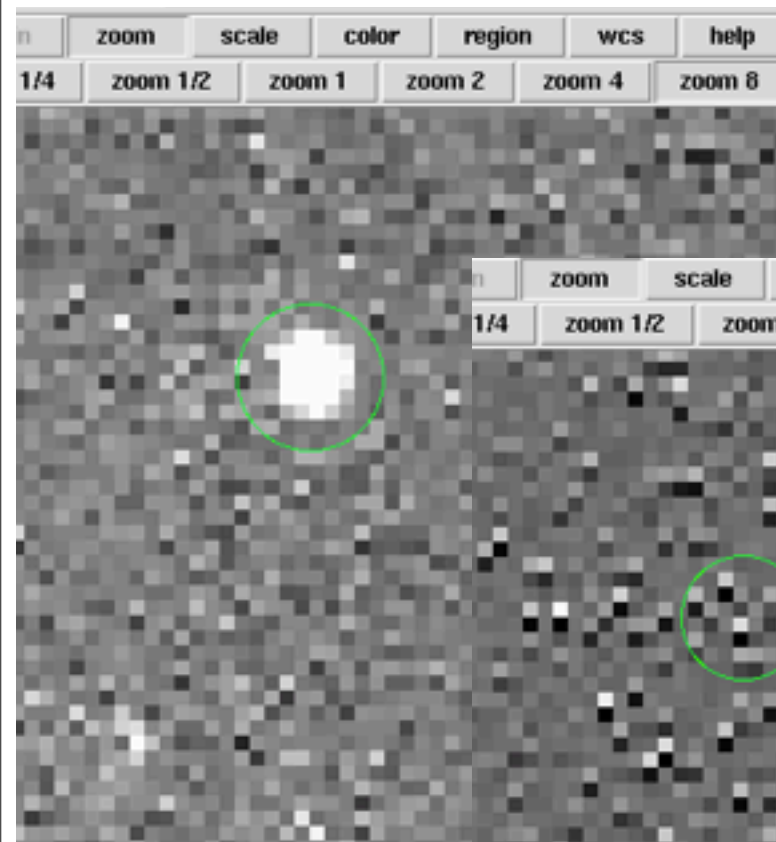
illum_ccmap-*.fits
same x,y
Top = 214.4
Bottom = 188.9

+/- 12.8 from mean
+/- 6.3% from mean

from original mean:
+6/-20
+2.9%/-9.6%

Dipoles reduced to
~background levels!

Dipoles at $x,y = (542, 1689)$ and $(543, 1686)$ --
above and below star center, so they have
opposite effects on profile

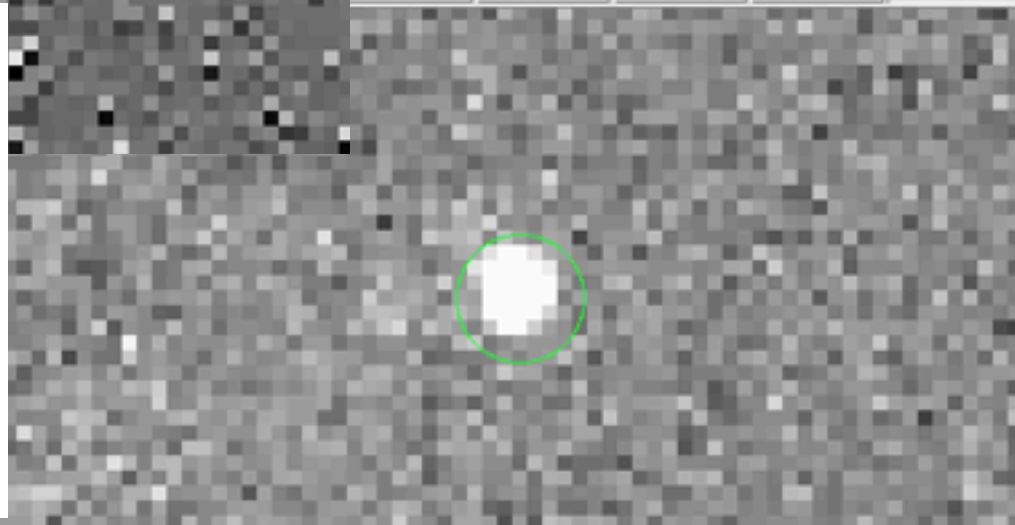


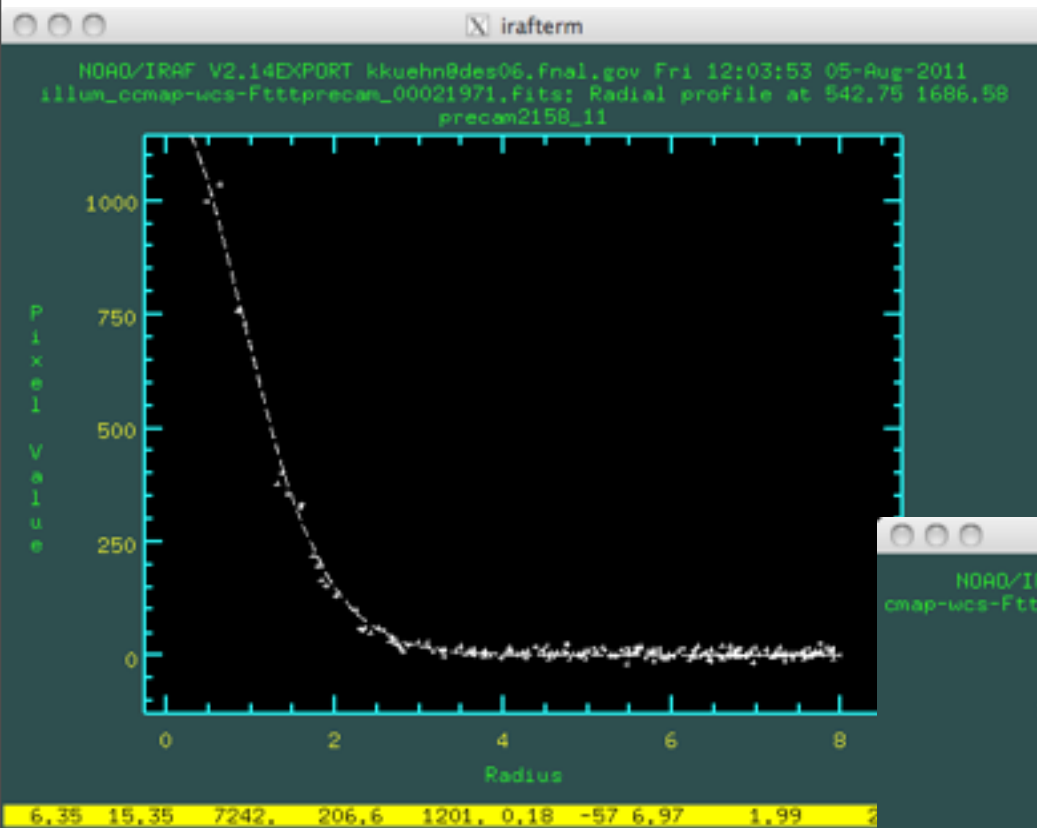
Before: Top pixel = 1327.8, Bottom pixel = 458.6
During: Top pixel = 1.069, Bottom pixel = 0.860
After: Top pixel = 1242.4, Bottom pixel = 533.1

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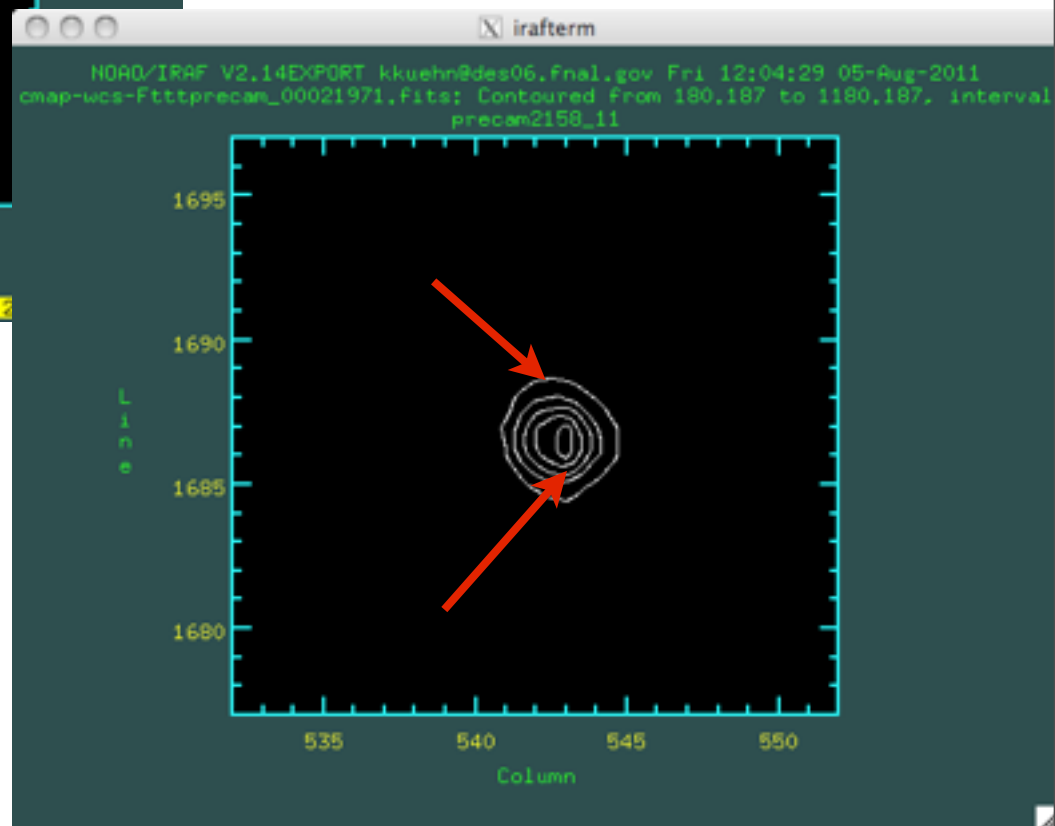
Before: Top pixel = 281.0, Bottom pixel = 471.7
During: Top pixel = 1.054, Bottom pixel = 0.877
After: Top pixel = 266.7, Bottom pixel = 538.0
Illum correction improves stellar profile, and

IMEXAM shows nearly perfect contours!





However, Illumination
correction didn't
improve photometry
significantly...

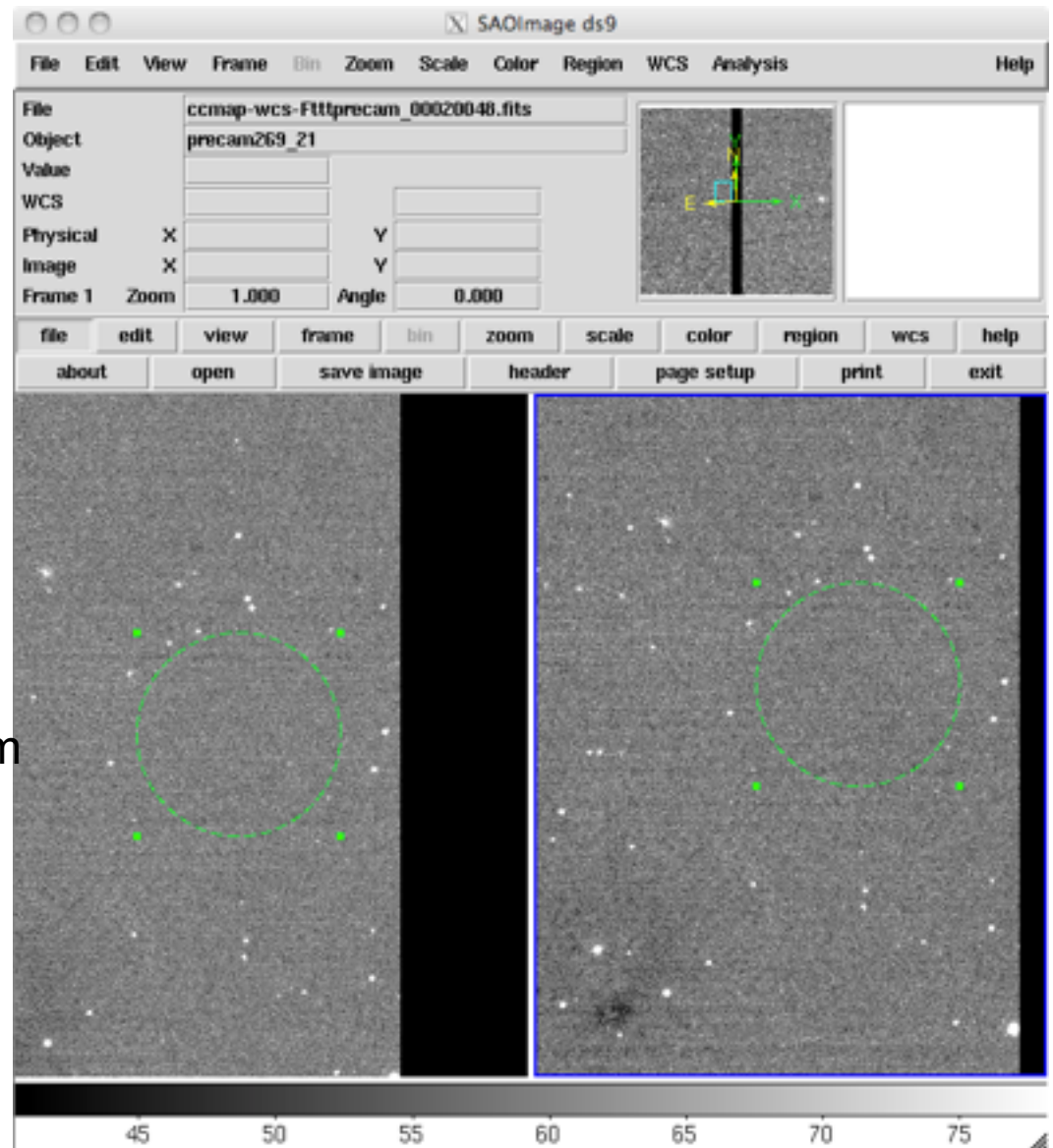


Imstat Tests:

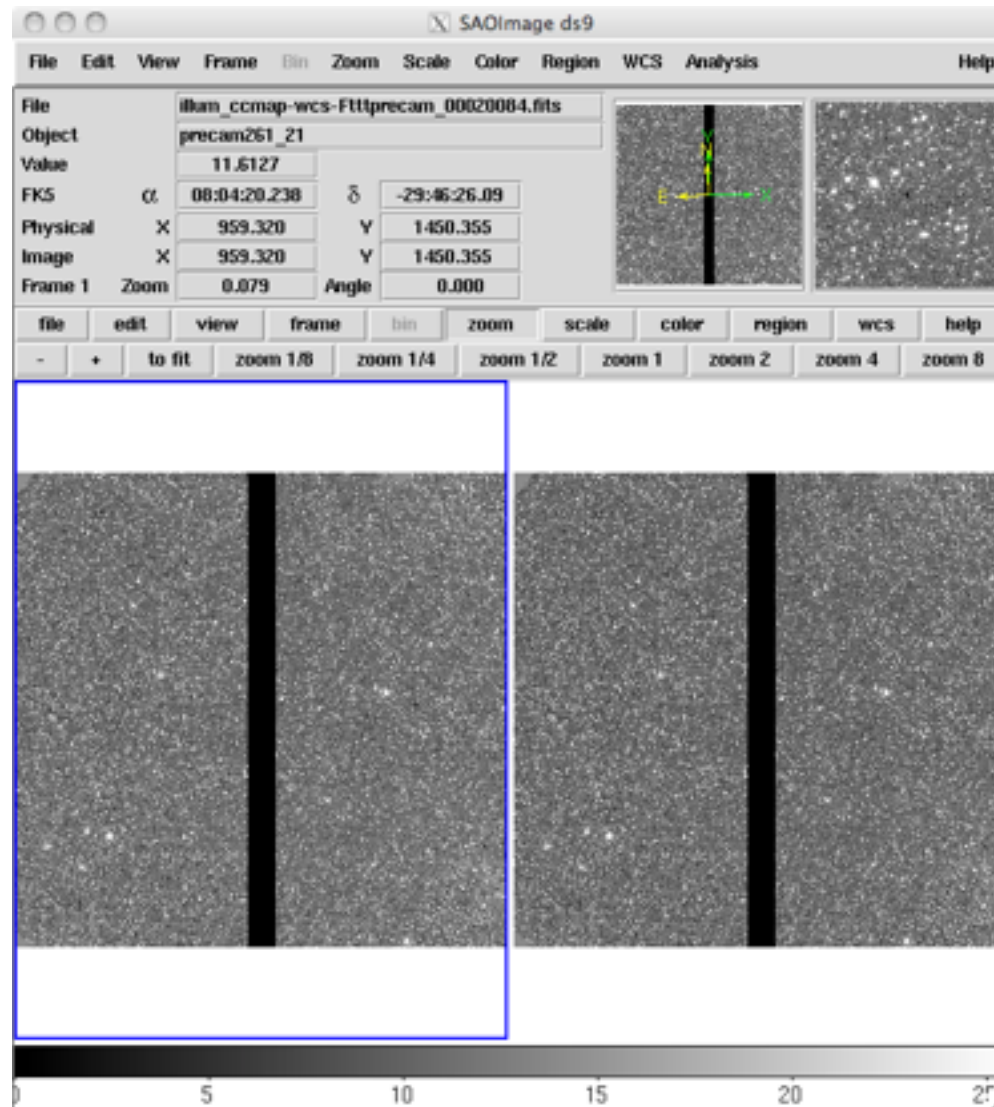
Before and after illum corr,
photometry was comparable, and
in some cases better without illum
correction.

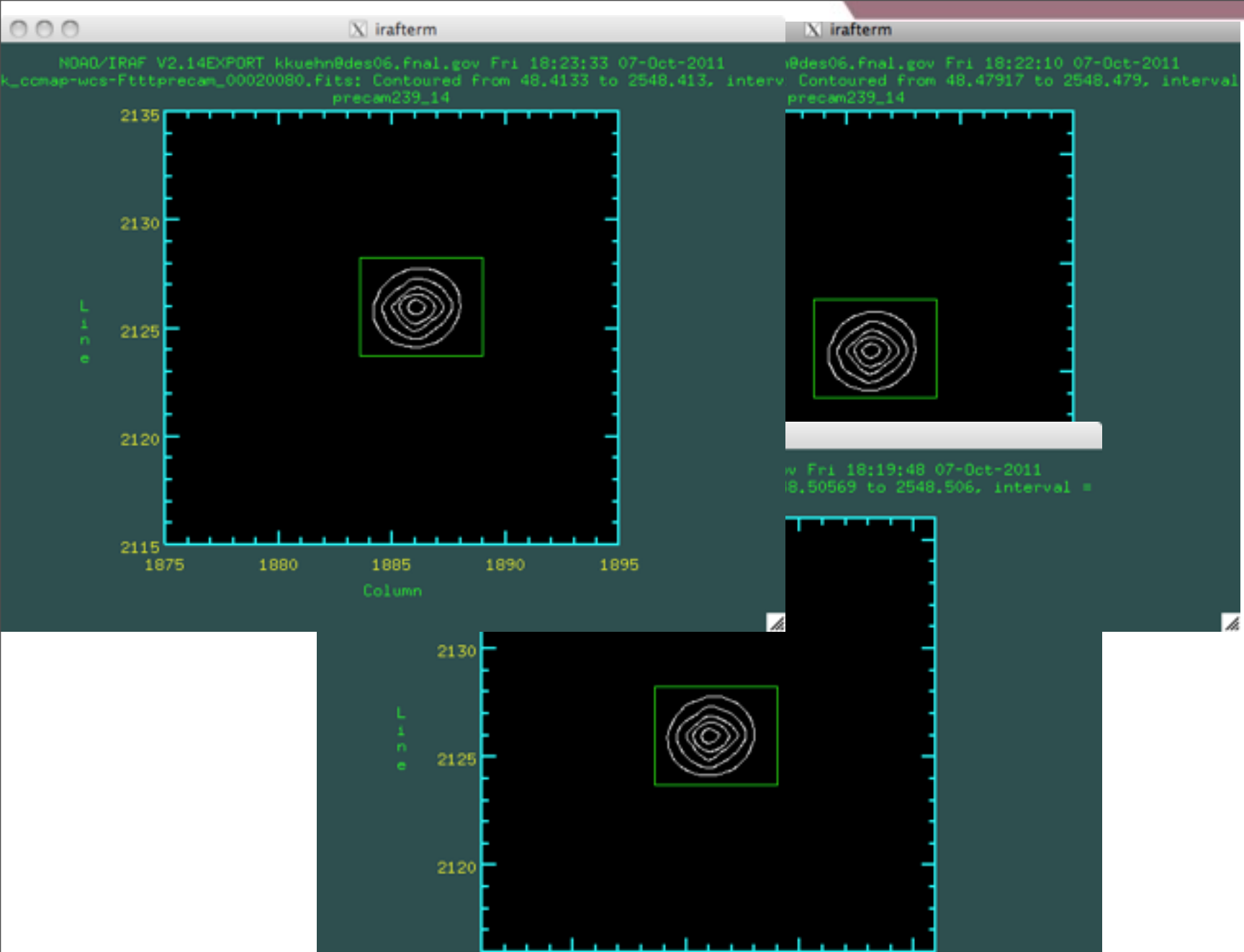
This was determined *not* to be an
effect of noise/statistics of images
that were combined to derive
illumination maps...

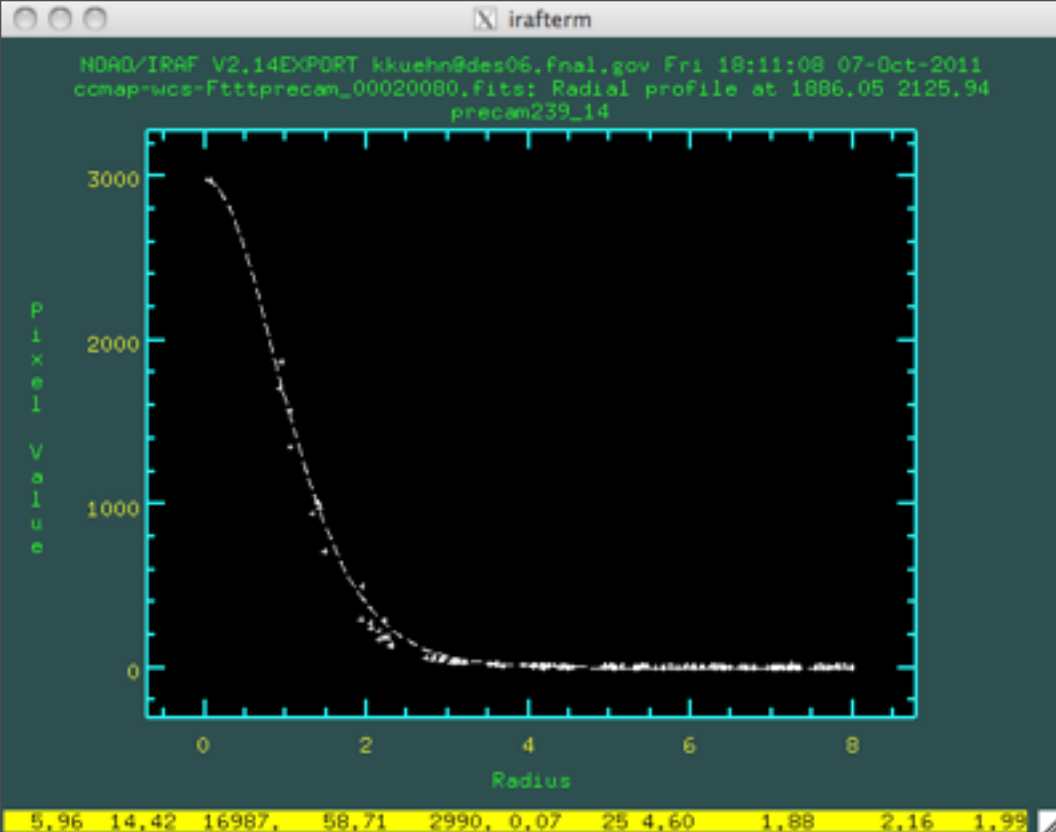
Comparable noise before and after
Illumination maps were created from
daily images as well as weekly...



Illumination map comparison 1 day vs. 1 week



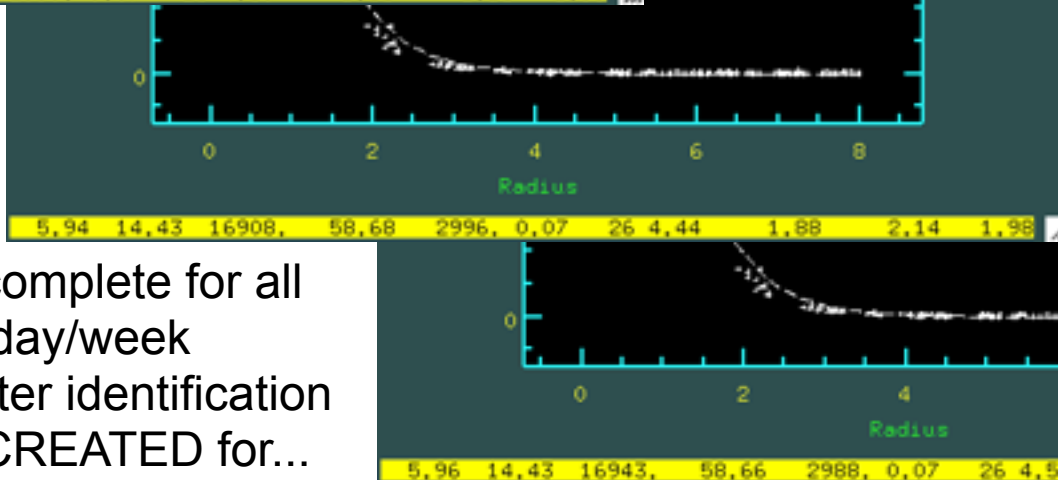




marginally different statistics for
ccmap, illum, illum_week...

v Fri 18:09:16 07-Oct-2011
ial profile at 1886.05 2125.94

i 18:12:46 07-Oct-2011
l profile at 1886.06 2125.94



Illum maps nearly complete for all
Golden Nights, for day/week
Includes stuck shutter identification
NEW .CAT FILES CREATED for...



v3 (SExtractor) Photometry

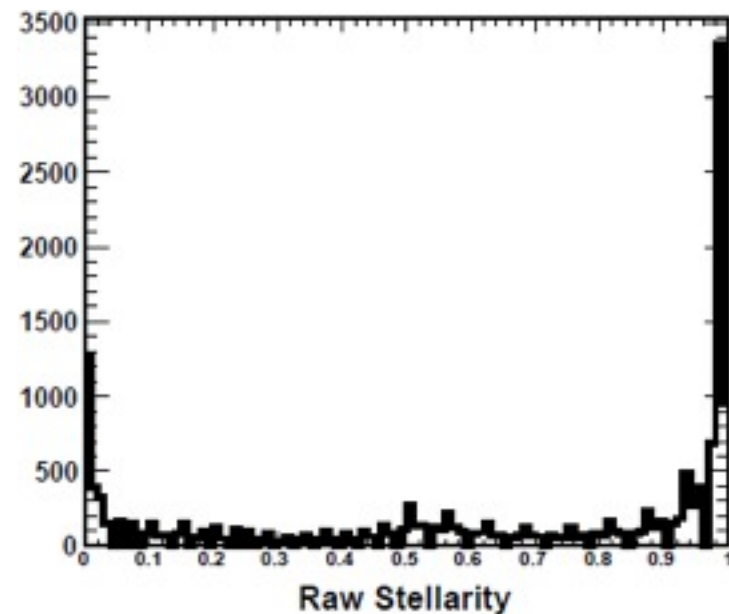
- Using FNAL v3 production + ANL ROOT scripts
- Applying SDSS airmass correction
- Using USNO, Southern u'g'r'i'z', and SDSS standards
- Cutting USNO, Southern, and SDSS $\text{magerr} < .01$ for all bands
- “v3aper10”
 - aper10 and $\text{fwhm} < 4$.
 - $\text{stellarity} > .95$
- Starflats applied to SDSS



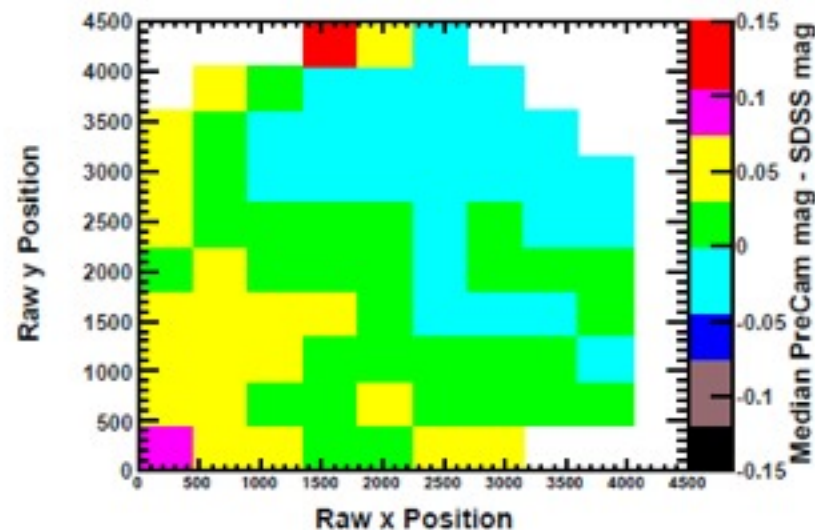
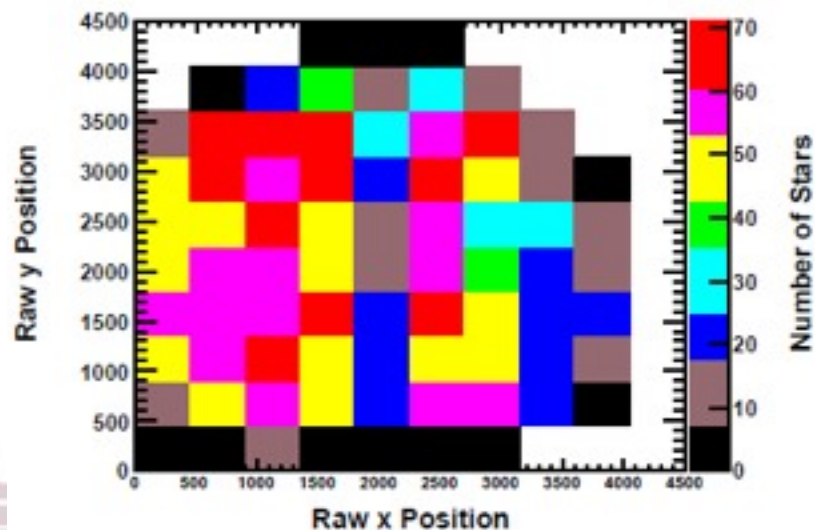
Evaluation of photometric error improvement per cut

Version	Standard Deviation
v2aper12	.06973
v3aper12	.06871
v3aper12, stellarity > .95	.04901
v3aper10, stellarity > .95	.048496
v3aper10, stellarity > .95, fwhm < 4.	.048434
v3aper10, stellarity > .95, fwhm < 4. , pixels cut	.048447
v3aper10, stellarity > .95, fwhm < 4. , pixels cut, starflats	.040106
v3aper10, stellarity > .95, fwhm < 4. , pixels cut, starflats, mag < 17.	.03838

20110108 SDSS r-band science images



Pixels cut: 2010<xpos<2250



v2aper12 vs. v3aper10 - Southern Standards

Date	Zero-Point Offset (Southern Standards) – v2, aper12	Zero-Point Offset (Southern Standards) – v3, aper10	Standard Deviation (Southern Standards) – v2, aper12	Standard Deviation (Southern Standards) – v3, aper10
20110107	g: .7747	g: .81565	g: .05225	g: .02219
20110112	g: .7955	g: .821	g: .02195	g: .02043
	r: .1805	r: .16541	r: .04771	r: .01888

- Using science images
- 3 outliers are from image 21882
- Image 21882 has diagonal streak in ds9 and is not flagged in bad image filters.



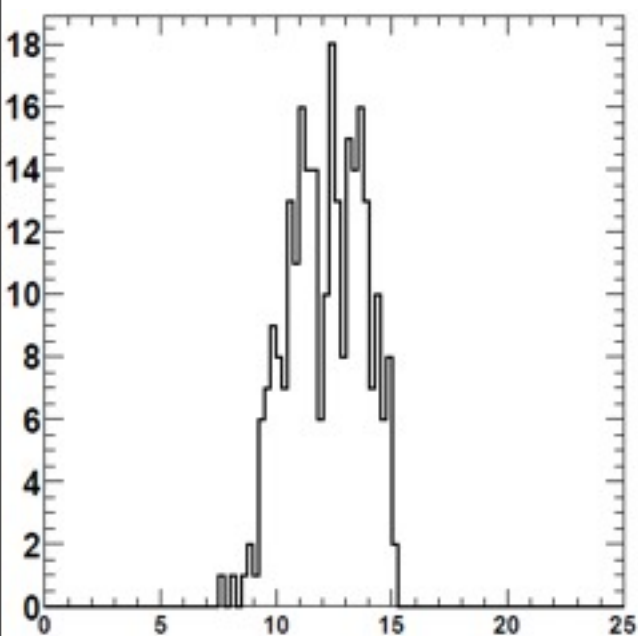
v2aper12 vs. v3aper10 - SDSS

Date	Zero-Point Offset (SDSS), v2 aper12	Zero-Point Offset (SDSS), v3 aper10	Standard Deviation (SDSS)-NoMagCut, v2 aper12	Standard Deviation (SDSS)-NoMagCut, v3 aper10	Standard Deviation (SDSS)- mag<17, v3 aper10
20110107	g: .83159 z: -.74	g: .81859 z: -.71	g: .08784 z: .07515 RA0to50	g: .04731 z: .0446 RA0to50	g: .04554 z: .0445 RA0to50
20110108	r: .21637 z: -.686	r: .22637 z: -.6552	r: .06973 z: .07761 RA0to50	r: .04011 z: .04484 RA0to50	r: .03838 z: .04484 RA0to50
20110112	i: .03895 z: -.7865	i: .03044 z: -.7645	i: .08617 z: .06198 RA40to50	i: .0485 z: .03803 RA40to50	i: .04832 z: .03803 RA40to50

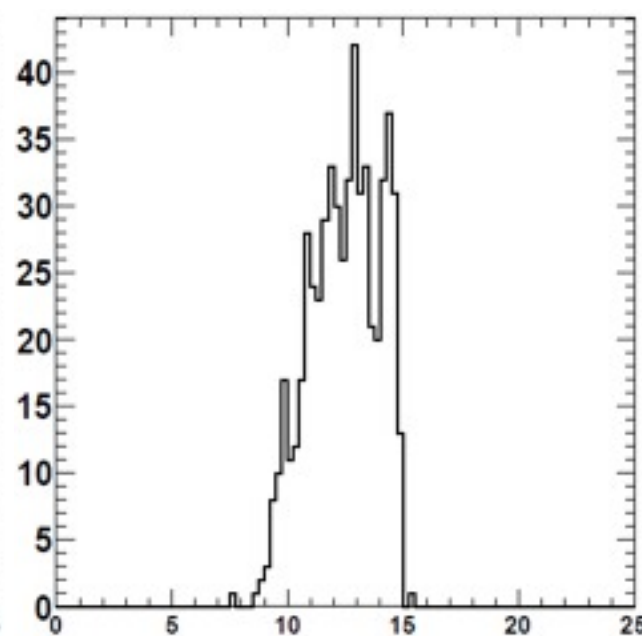
- Using science images
- Starflats have been applied



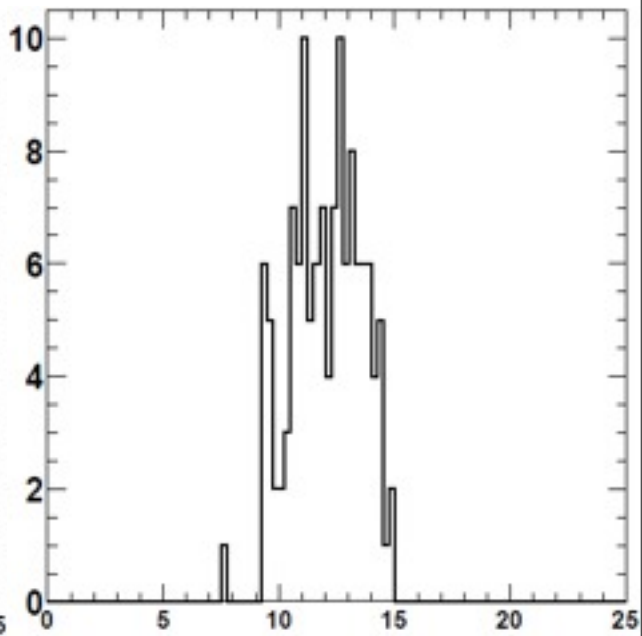
Standard Magnitudes, USNO



USNO Magnitudes
20110107



USNO Magnitudes
20110112

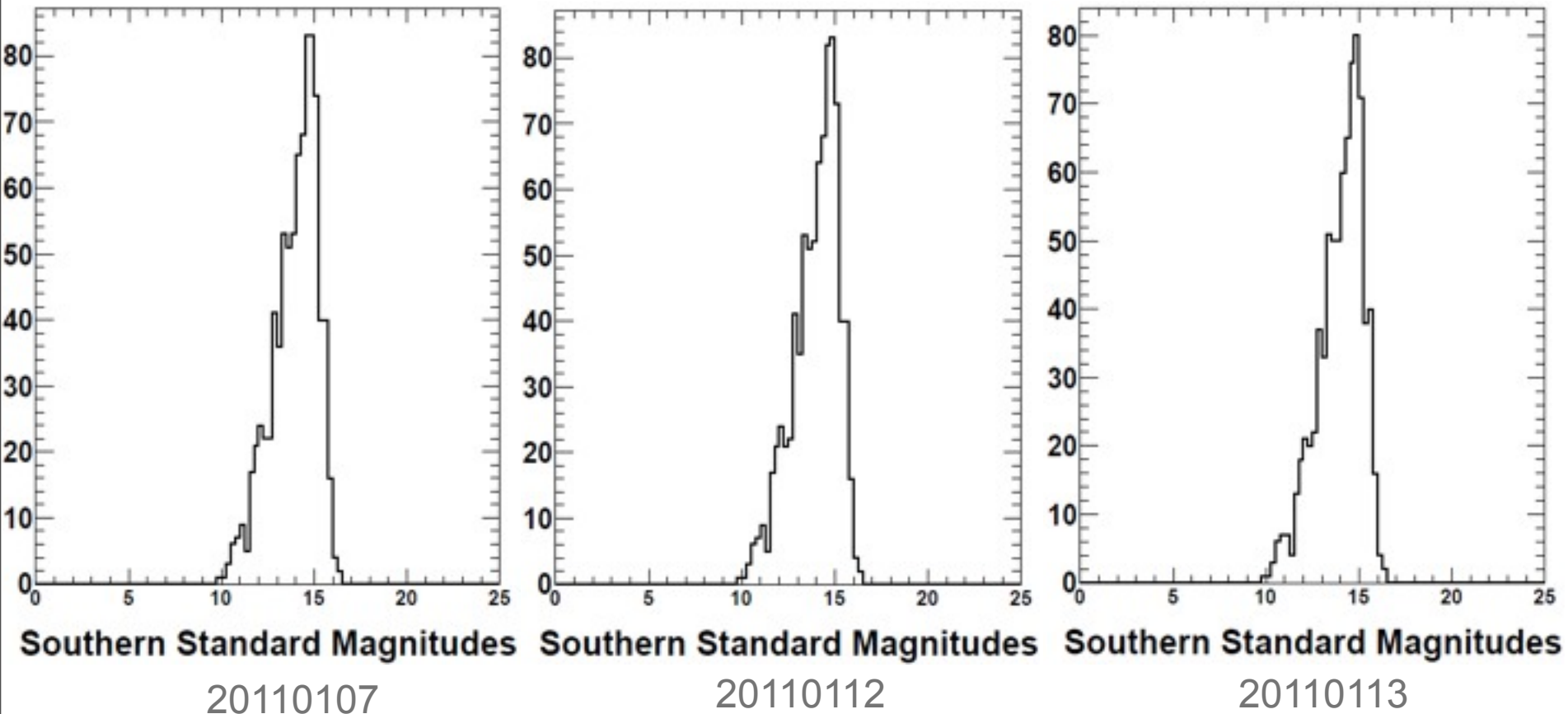


USNO Magnitudes
20110113

- All plots have $\text{magerr} < .01$



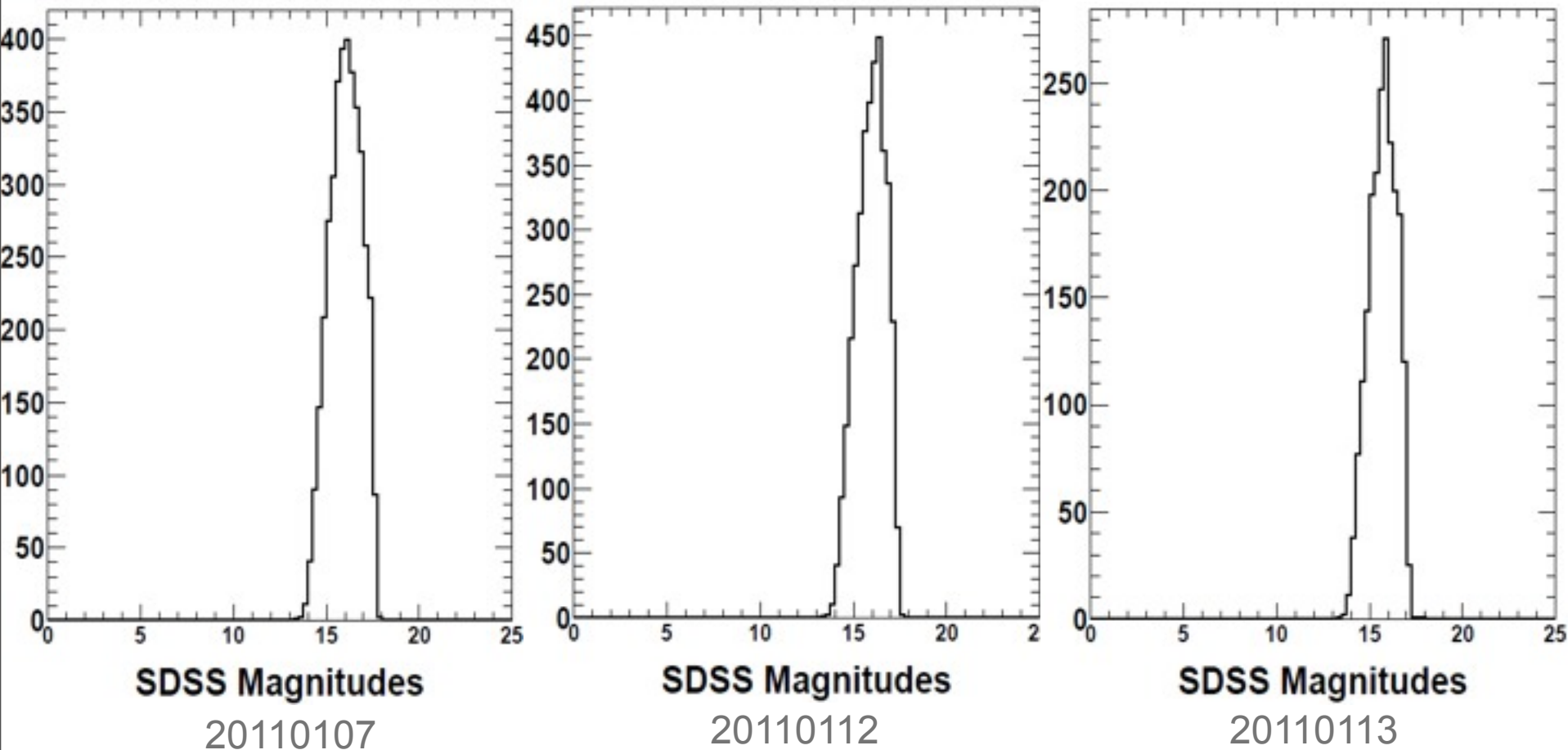
Standard Magnitudes, Southern Standards



- All plots have $\text{magerr} < .01$



Standard Magnitudes, SDSS



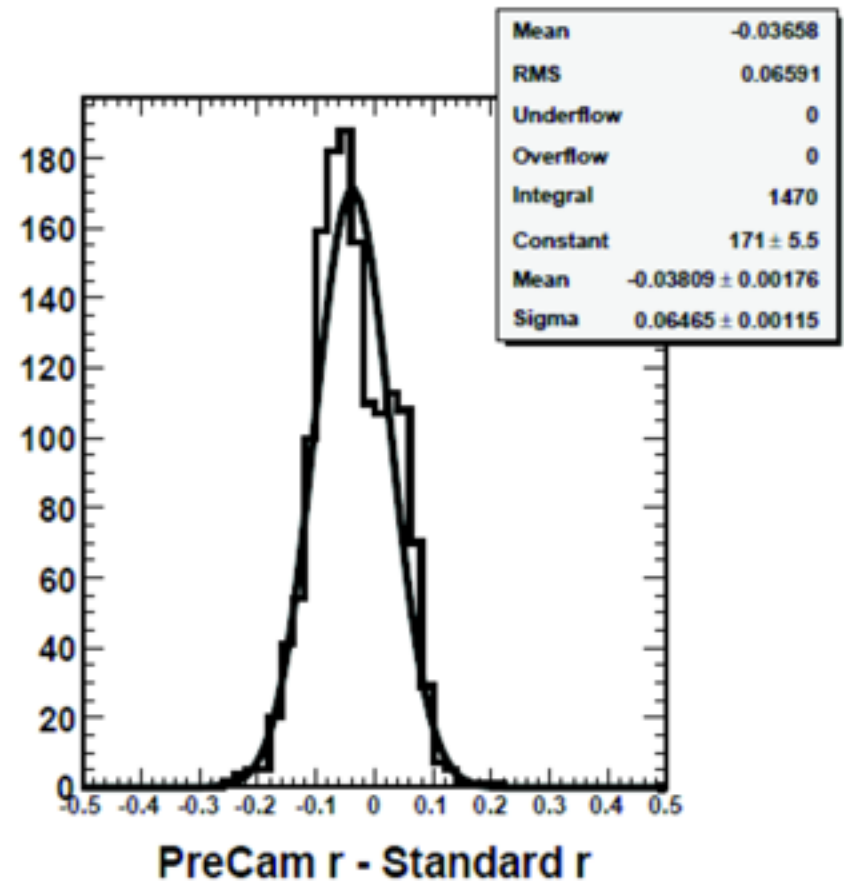
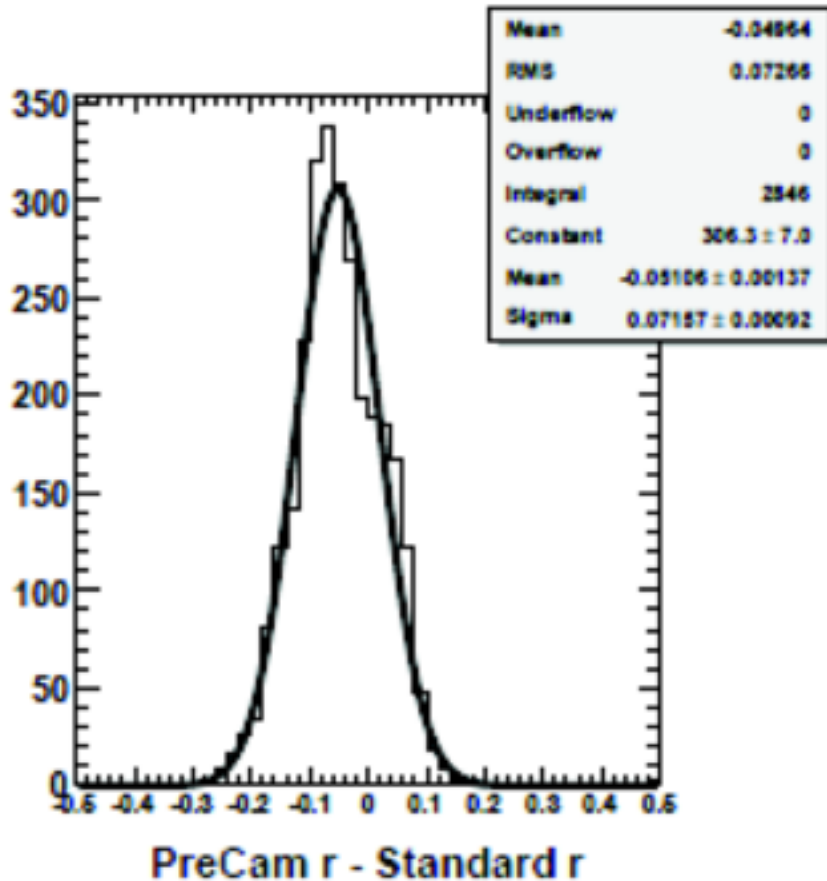
- All plots have $\text{magerr} < .01$



Date	Zero-Point Offset (USNO)	Standard Deviation (USNO)	Zero-Point Offset (Southern Standards)	Standard Deviation (Southern Standards)	Zero-Point Offset (SDSS)- RA40to50	Standard Deviation (SDSS)- mag<15	Standard Deviation (SDSS)-NoMagCut
20101215			g: 2.3372	g: .05323	g: 2.17425	g: .05258	g: .08925
			r: 2.1269	r: .05176	r: 1.95941	r: .04194	r: .09445
			i: 2.2864	i: .05021	i: 2.12993	i: .0581	i: .1065
			z: 2.5072	z: .06227	z: 2.32859	z: .05982	z: .08172
20110107	g: 2.08978	g: .02374	g: 2.0802	g: .02843	g: 2.0645	g: .04711	g: .09277
	r: 1.899857	r: .03213	r: 1.91872	r: .04228	r: 1.9346	r: .04305	r: .09258
	i: 2.05227	i: .03222	i: 2.05298	i: .02944	i: 2.07891	i: .05134	i: .09468
	z: 2.247	z: .02319	z: 2.263	z: .04915	z: 2.2993	z: .05304	z: .08514
20110108	g: 2.1784	g: .07305	g: 2.12746	g: .03003	g: 2.1617	g: .05267	g: .1057
	r: 1.98041	r: .06301	r: 1.93154	r: .04221	r: 1.94502	r: .03744	r: .09614
	i: 2.1281	i: .05031	i: 2.24	i: .04874	i: 2.10669	i: .05214	i: .1025
	z: 2.3587	z: .05476	z: 2.3044	z: .03008	z: 2.3456	z: .05843	z: .07865
20110112	g: 2.1035	g: .03165	g: 2.07424	g: .02947	g: 2.11098	g: .04387	g: .08868
	r: 1.932	r: .05489	r: 1.905615	r: .03518	r: 1.92643	r: .03939	r: .08243
	i: 2.0765	i: .04316	i: 2.06179	i: .03624	i: 2.07017	i: .04554	i: .102
	z: 2.248	z: .04514	z: 2.21012	z: .03695	z: 2.25469	z: .05621	z: .08538
20110113	g: 2.08618	g: .02186	g: 2.07	g: .03127	g: 2.143606	g: .02575	g: .09088
	r: 1.90392	r: .02544	r: 1.89748	r: .03662	r: 1.9298	r: .04268	r: .08401
	i: 2.05038	i: .02691	i: 2.06527	i: .04353	i:	i:	i:
	z: 2.21058	z: .02033	z: 2.20684	z: .03638	z: 2.26745	z: .06571	z: .0936

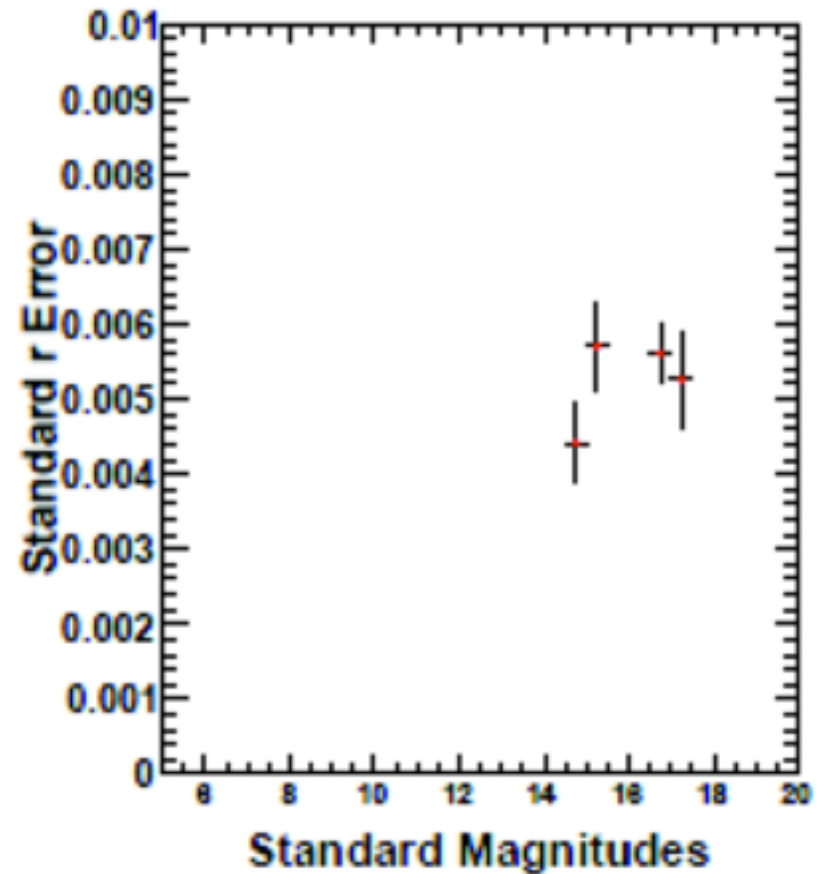
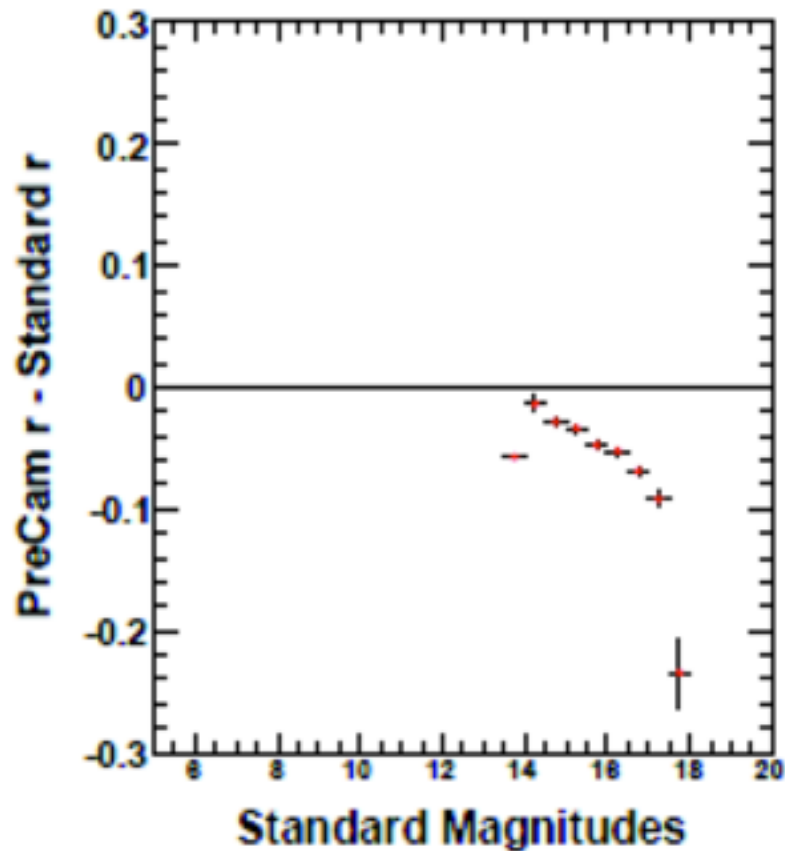
Monday, October 10, 2011

DAOPHOT Results



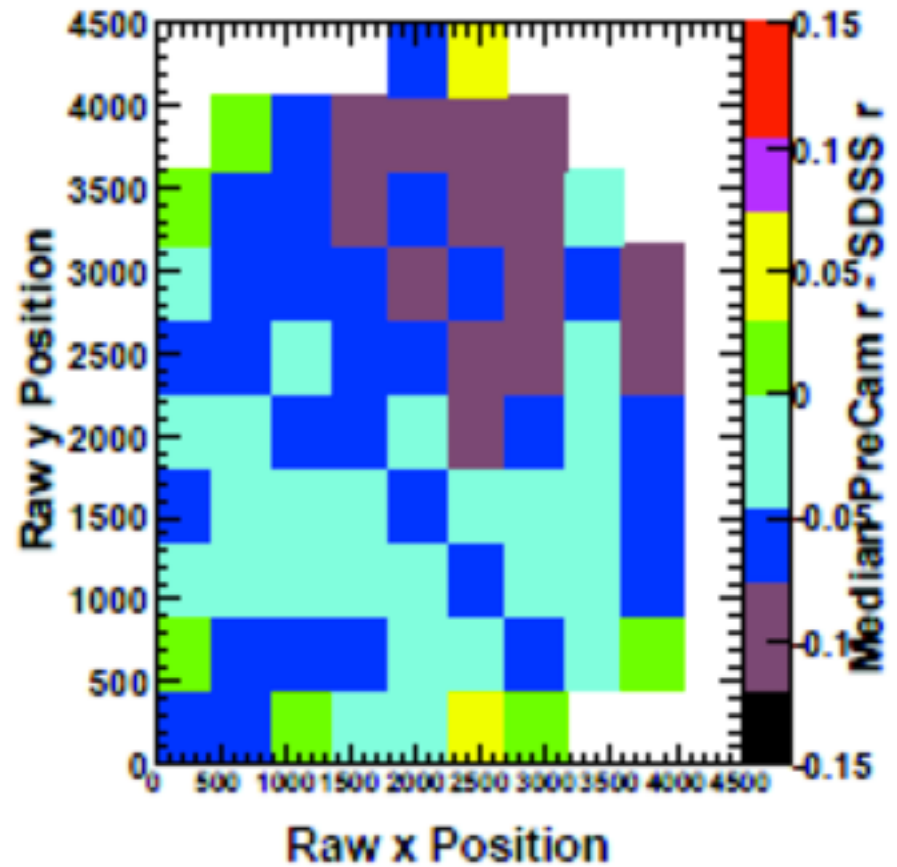
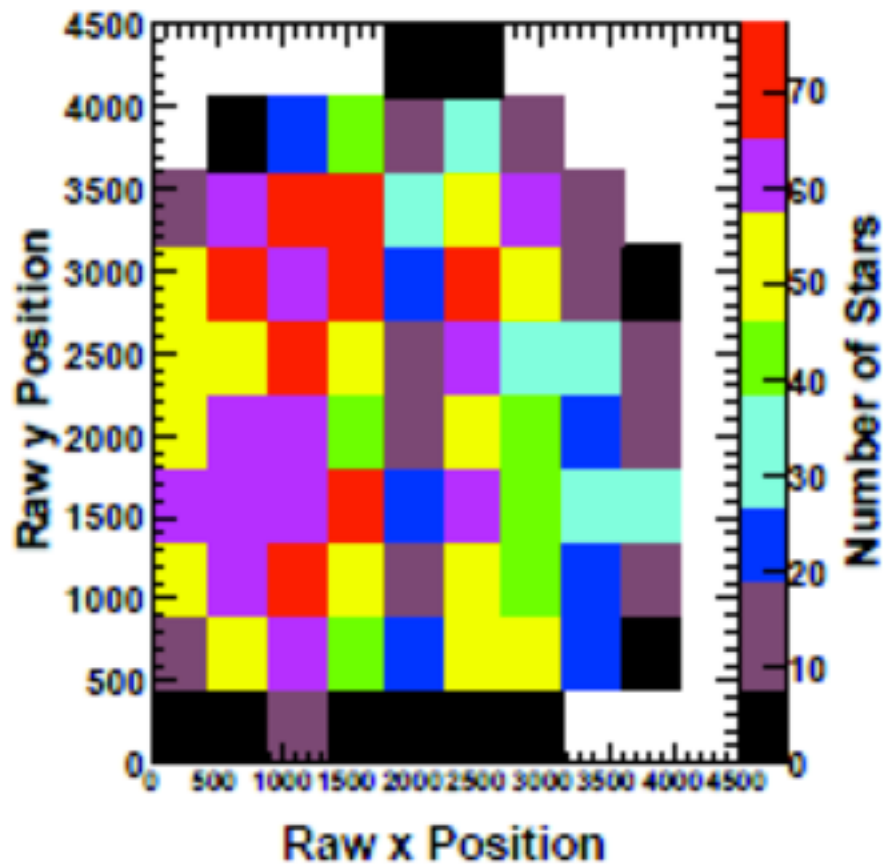
- Six images from 8 Jan were analyzed with SDSS matches.
 - With just Douglas' cuts ($\text{FWHM} < 4$, $\text{stellarity} > 0.95$), $\sigma = 0.07$ (left).
 - With an additional $\text{magcut} < 16$, $\sigma = 0.065$ (right).

DAOPHOT Results (cont.)



- Results from Douglas' cuts only (no additional mag cut)
- Systematic offset? Is DAOPHOT configured properly?

DAOPHOT Results (cont.)



- Results from Douglas' cuts only (no additional mag cut).
- Error Correlated with position...?

To Do:

- Put everything together...v3+ processing, bad image removal & illumination correction, then:
- Photometry
 - DAOPHOT?
 - PSFEx
 - SDSS reCalibration? (Errors of $\sim 4\%$ > USNO, Southern Standard of 2%)
- Catalog
- Incorporate Streak Correction, Bad Image Detection/Correction, Illumination Maps into Brazil Pipeline

